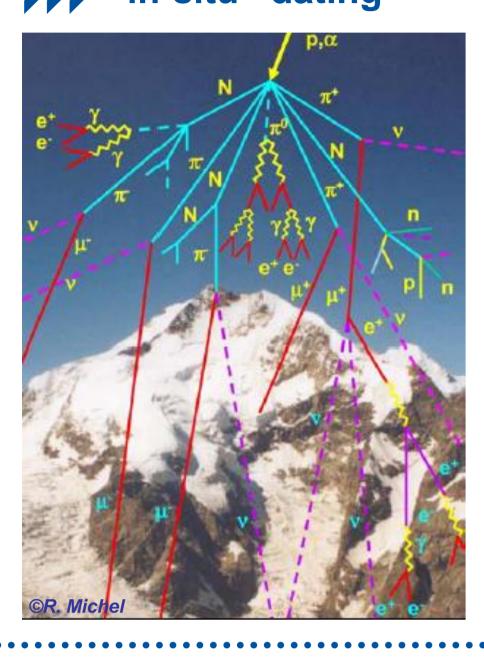
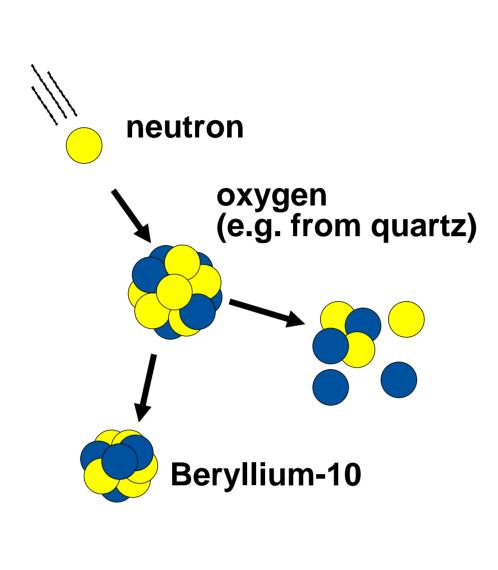
How to get a new accelerator mass spectrometry (AMS) facility running: The chemistry part

S. Merchel¹, M. Arnold², G. Aumaître², D. Bourlès², R. Braucher²

Cosmic radiation on Earth

- \blacksquare MeV/GeV particles (p, α , heavy ions)
- shielding (flux & energy) and transformation e.g. into neutrons by the Earth's magnetic field & atmosphere
- elemental conversion by nuclear reactions in the atmosphere (14N(n,p)14C) >>>> radiocarbon-dating e.g. for archaeology or climate reconstruction from ice cores
- elemental conversion by nuclear reactions in terrestrial materials / rocks (so-called "in-situ"-production) >>> with time concentration of (radio-) nuclides increases
- irradiation of a "fresh" surface (e.g. after a volcanic eruption), reconstruction of "starting time" possible"in-situ"-dating



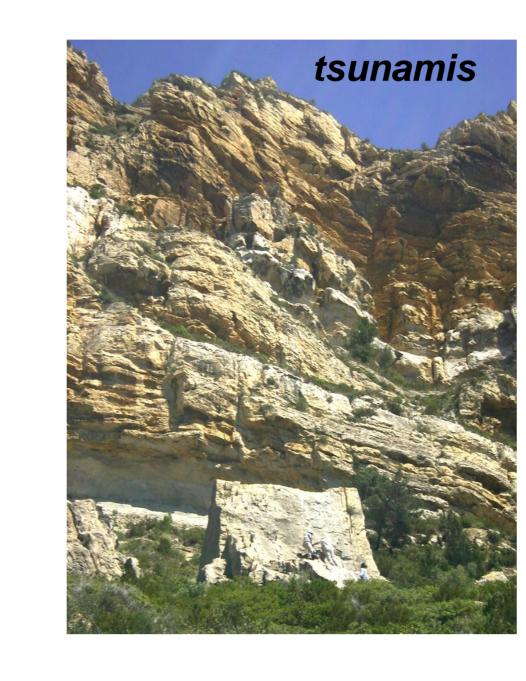


"In-situ"-produced cosmogenic nuclides (CN)

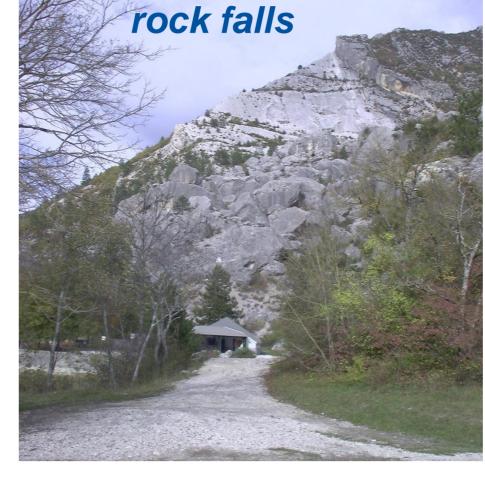
everything having a "fresh" surface can be dated

volcanic eruptions











Indirect dating by CN

climate change, e.g. via glacier movements



Special case (irradiation in space): Extraterrestrial material

- transfer times from the meteorite's parent body (asteroids, Moon, Mars)
 irradiation age
- residence time at the place of discovery (e.g. hot desert, Antarctica) before somebody takes the meteorite home

>>> terrestrial age





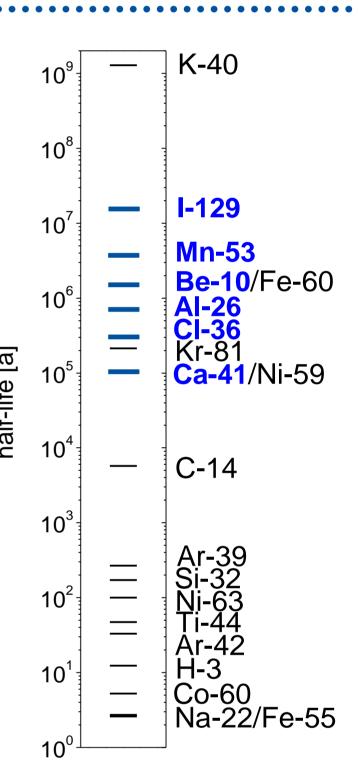
Sample preparation for accelerator mass spectrometry (AMS)

Definition of "good" AMS-chemistry

- "good " standards, i.e.
 - traceable
 - different ratios for volatile elements to avoid cross-contamination (³⁶Cl, ¹²⁹l)
- **■** low machine blanks (¹0Be)
- low stable nuclide carrier for chemistry blanks (10Be)
- real sample chemistry for different matrices & nuclides
 - fast and low-cost (chemicals, man power, lab space)
 - safe (also for non-chemists)
 - low risk of cross-contamination & contamination
 - producing "pure" targets, i.e. high stable nuclide current & low isobar concentration
 - >>> good statistics, detection limits & high sample throughput

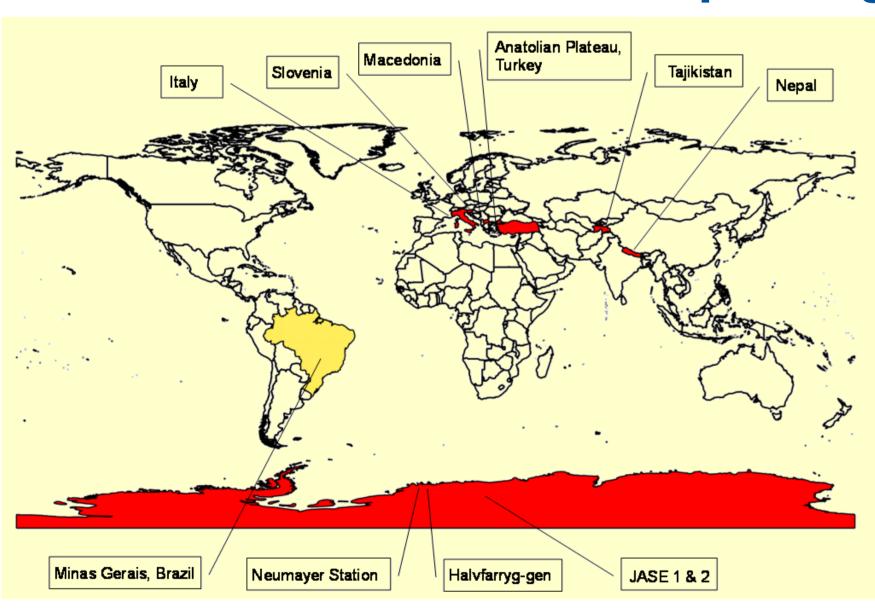
Status @ DREAMS

- standards for ¹⁰Be, ²⁶AI, ³⁶CI, ⁴¹Ca, ¹²⁹I traceable via round-robins or cross-calibration elsewhere [ARN10, MER04, MER09, MER11]
 - Special case ¹⁰Be: produced via ⁰Be(n_{th},γ)¹⁰Be @ TRIGA, Atominstitut Vienna >>> (1.73±0.02)⋅10⁻¹² ¹⁰Be/⁰Be
 - ²⁶AI, ³⁶CI, ⁴¹Ca & ¹²⁹I @ 3 different ratios each 10⁻⁹ 10⁻¹³
- "home-made" ¹⁰Be carrier & machine blank from shielded Be-containing mineral (phenakite Be₂SiO₄) [MER08]
- speeded-up ¹ºBe-chemistry for ice core sample (5 d → 24 h
 / 10 samples)
- development of ⁵³Mn-chemistry for "in-situ" samples (marcasite/pyrite/realgar)
- ¹¹ºBe & ³6Cl targets from "in-situ" samples (SiO₂ & CaCO₃) measured @ ASTER & VERA >>>> chemistry blank at least one order of magnitude lower than samples & high current
- chemistry training of external partners (non-chemists) for ¹⁰Be & ³⁶Cl from "in-situ" samples satellite labs @ U Rennes & U Freiberg





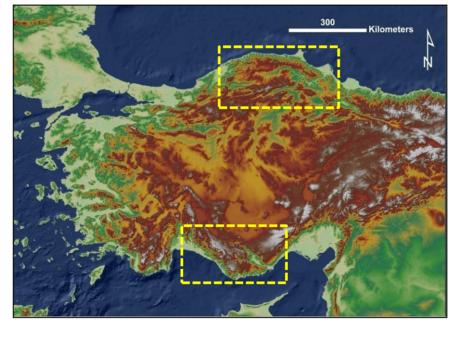
Selection of sample origins & applications



Historical rock fall @ Veliki Vrh, Slovenia
 Calibration site for production rate (36Cl)

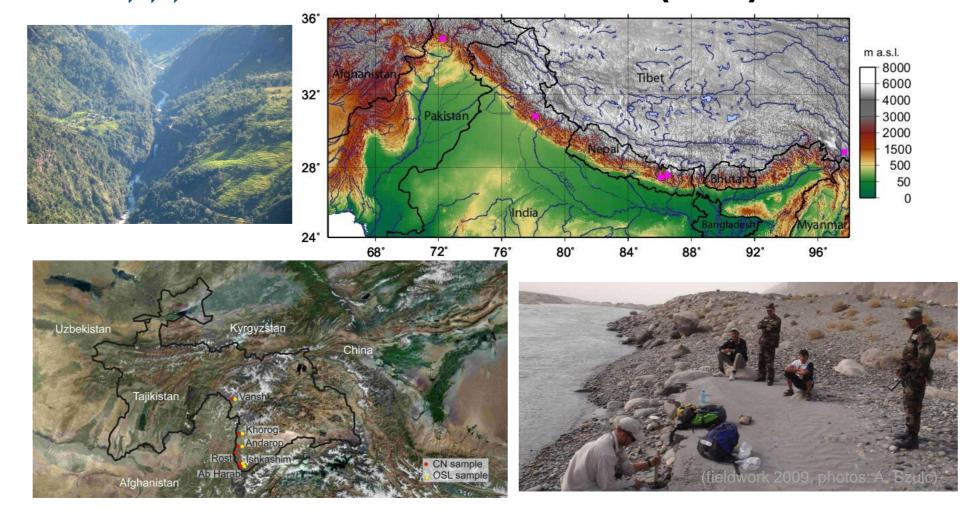


Vertical Anatolian Movement Project
>>> Uplift rates (36CI)





Geomorphology @ Himalaya & Pamir
 Basin-wide erosion rates (10Be)



Summary & Outlook

- ▲ AMS chemistry are successfully running for routine sample preparation (¹ºBe, ²ºAI, ³ºCI, ⁴¹Ca, ⁵³Mn) & satellite labs for pretreatment of "in-situ" samples are installed
- machine blanks and standards are prepared, measured elsewhere, and partially (¹⁰Be, ²⁶Al, ³⁶Cl) measured @ DREAMS [AKH10]
- ⁴¹Ca & ¹²⁹I will be measured @ DREAMS this week
- sophisticated multi-nuclide chemistry for extraterrestrial material [MER99] such as Martian meteorites will start next week

Acknowledgements

We are grateful to E. Strub for providing 26 Al-activity, to A. Wallner & M. Bichler for performing the neutron-irradiation of 9 Be, to C. Varajão for providing Be $_{2}$ SiO $_{4}$ crystals, and to R.C. Finkel, L. Benedetti, W. Möller, HVEE, the FZD-operator- & AMS-team for great cooperation. The VERA-team really helped us with 10 Be cross-calibration and 36 Cl "in-situ" sample AMS-measurements (esp. P. Steier, A. Wallner). Finally, we would like to thank our first chemistry lab and DREAMS users for their "braveness" working with us right from the start, i.e. C. Andermann, M. Fuchs, R. Gloaguen, K Klemm (U Freiberg & U Rennes), T.F. Schildgen, C. Yildirim (GFZ & U Potsdam), D. Wagenbach (U Heidelberg), D. Fritzsche, H. Miller, T. Opel (Alfred-Wegener-Institut, Bremerhaven & Potsdam), I. Mrak (U Ljubljani).

References

[AKH10] S. Akhmadaliev et al., this meeting, Thursday, 17:30 h.

[ARN10] M. Arnold et al., Nucl. Instr. Meth. Phys. Res. B 268 (2010) 1954.

[MER99] S. Merchel & U. Herpers, Radiochim. Acta 84 (1999) 215.

[MER04] S. Merchel & W. Bremser, Nucl. Instr. Meth. Phys. Res. B 223-224 (2004) 393. [MER08] S. Merchel et al., Nucl. Instr. Meth. Phys. Res. B, 266 (2008) 4921.

[MER08] S. Merchel et al., Nucl. Instr. Meth. Phys. Res. B. 266 (2008) 4921. [MER09] S. Merchel et al., Geochim. Cosmochim. Acta (2009) A871.

[MER09] S. Merchel et al., Geochini. Cosmochini. Acta (2009) A67 1.

[MER11] S. Merchel et al., to be submitted to Nucl. Instr. Meth. Phys. Res. B